



IACSAT-1

The first space observatory of the IAC

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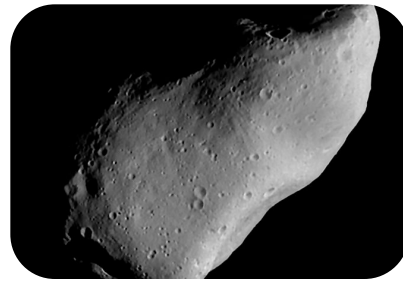
David Rodríguez



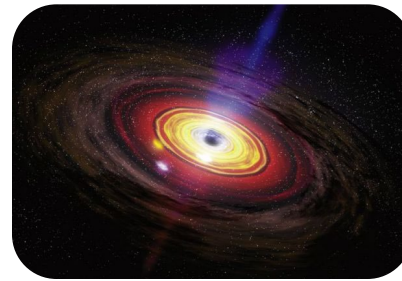
IACSAT-1: Multi-purpose space observatory



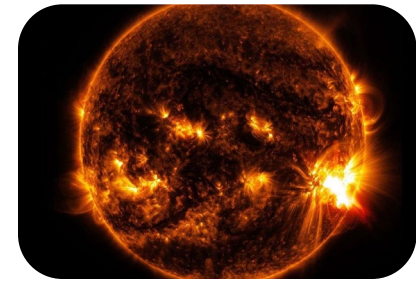
Exoplanets



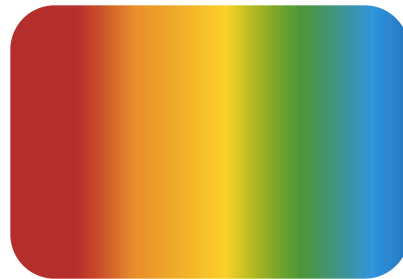
Primitive
asteroids



Blazars



Stellar activity



VIS-NIR
photometry



NUV
spectroscopy

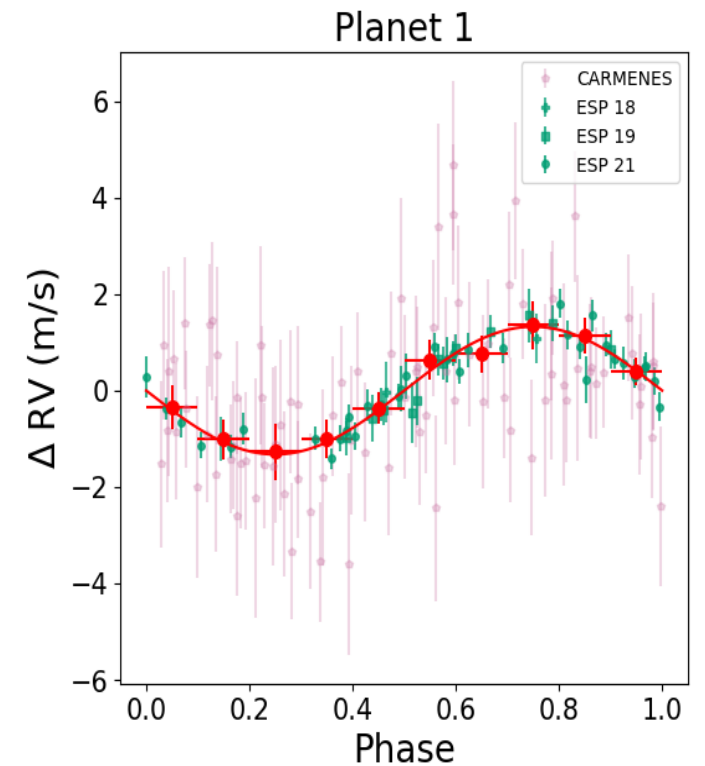
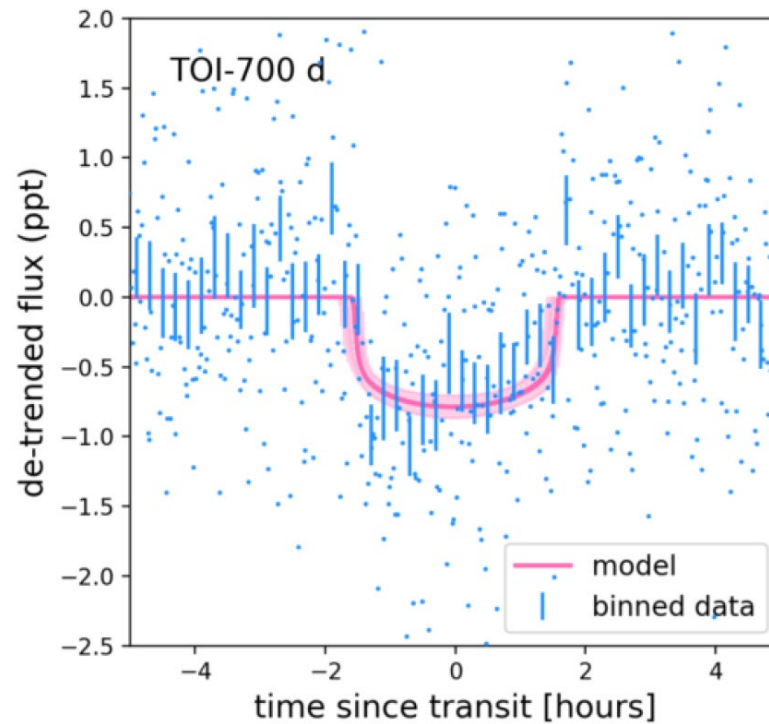
Confirmation of Earth-like exoplanet candidates

Earth-like candidates from TESS are challenging to validate using ground-based telescopes

Typical characteristics:

- M-dwarfs ($< 0.5 M_{\text{sun}}$)
- Vmag 11-14
- Radius < 2 Earth radii
- Insolation < 5 Earth flux

Prime targets for mass measurements and atmosphere characterization to search for biomarkers.



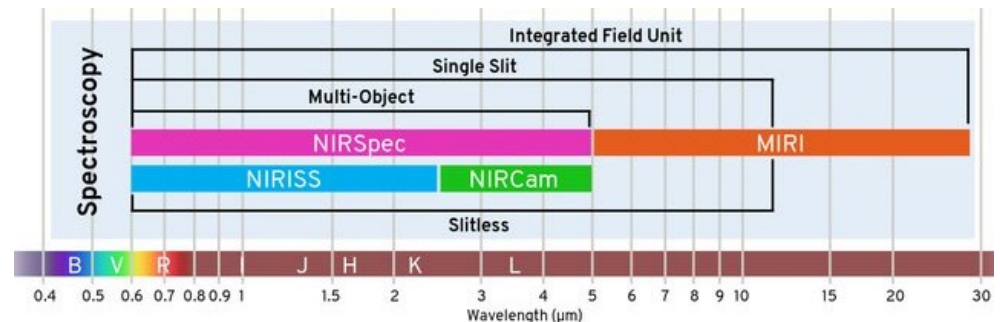
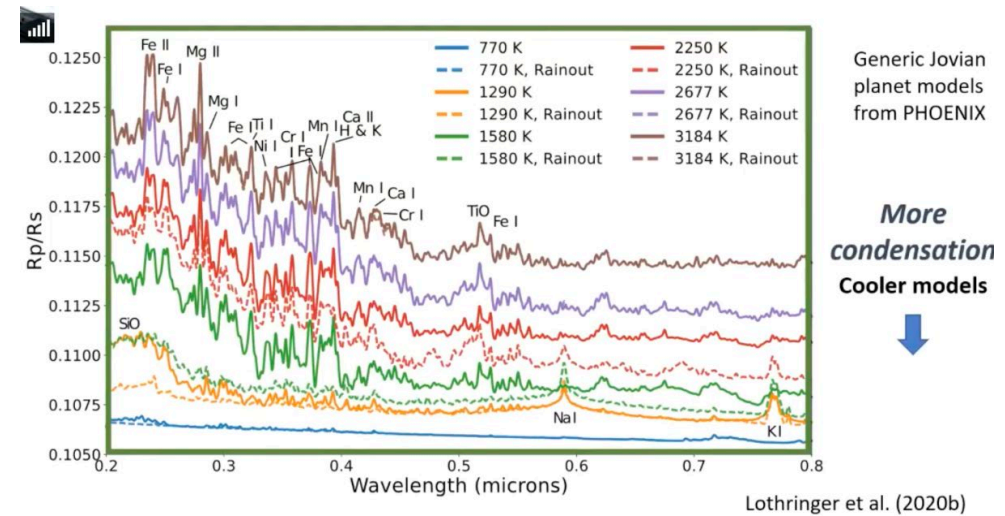
Exoplanet atmospheres

Rayleigh scattering by atmospheric hazes dominates the NUV region

Measure radii differences with wavelength in the NUV region

Distinguish between atmospheric metallicity and height of the cloud deck location

Completes the transmission spectra expected with JWST and ground-based facilities



Primitive asteroids

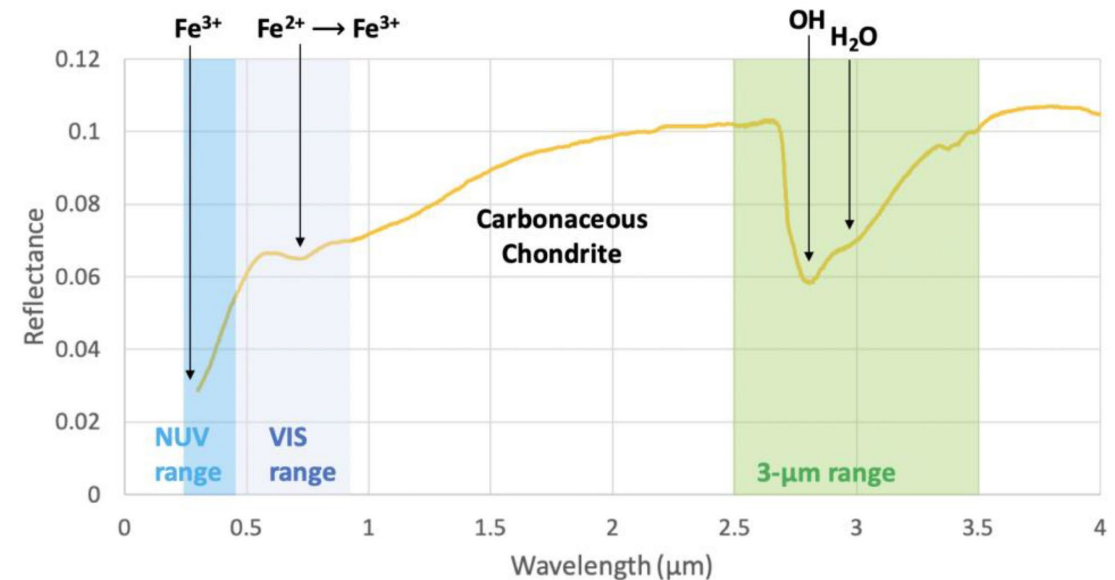
Shallow absorption in VIS range

Deep absorption in 3-microns

Deep absorption in NUV range

- Not typically explored

Correlations between UV, 0.7- μm , and 3- μm absorptions could help us understand the asteroids composition



Characterization of blazars

Blazars show two broad bands peaking at gamma rays and NUV regions

Disentangle mechanism behind SED.

- Leptonic vs pure hadronic emission models

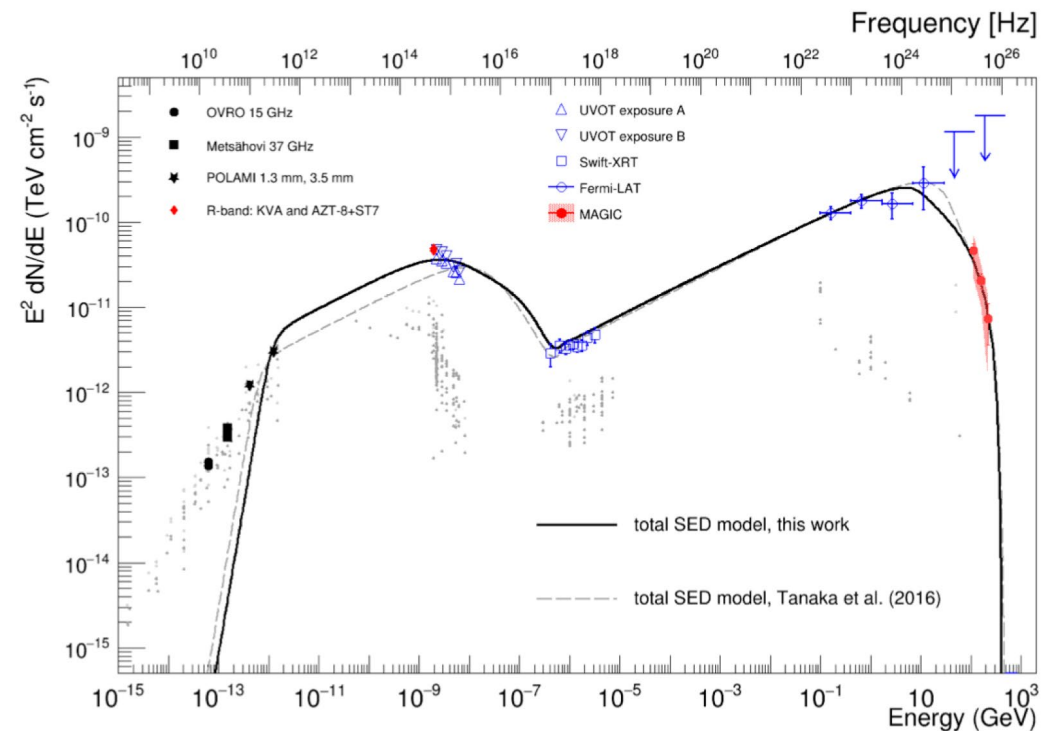
Leptonic framework

- Predicts correlation

Hadronic framework

- Predicts no correlation

Observations simultaneous with Fermi-Lat and Cherenkov telescopes

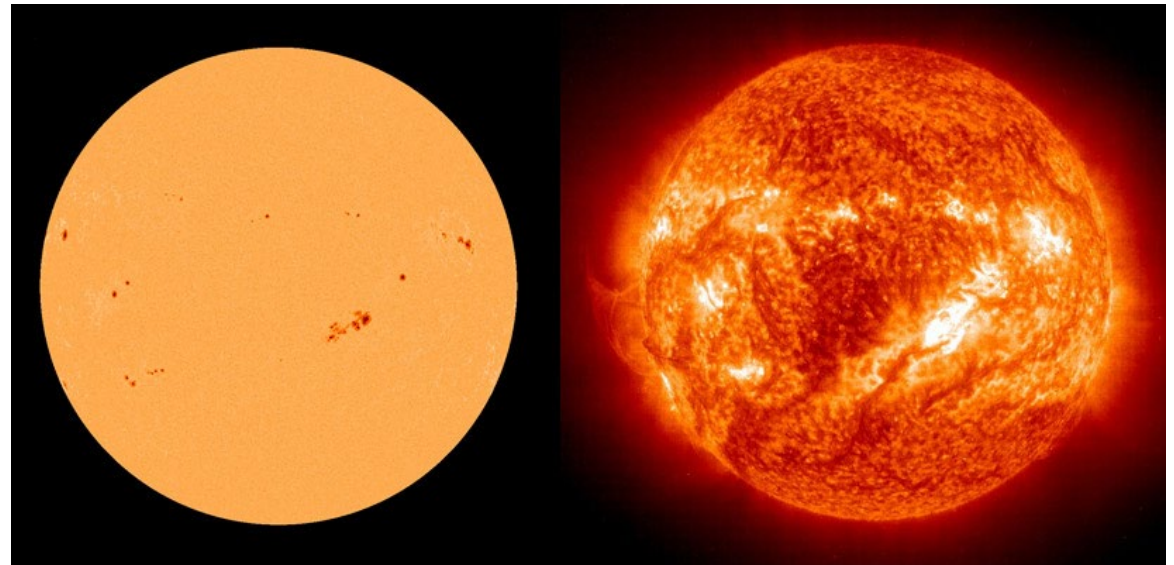


Stellar activity

Different active regions visible at different wavelengths

Contemporaneous VIS+NUV to disentangle the nature of the observed variations

Study a sample of stars to put constraints on the dynamo models



IACSAT-1

22 cm primary mirror

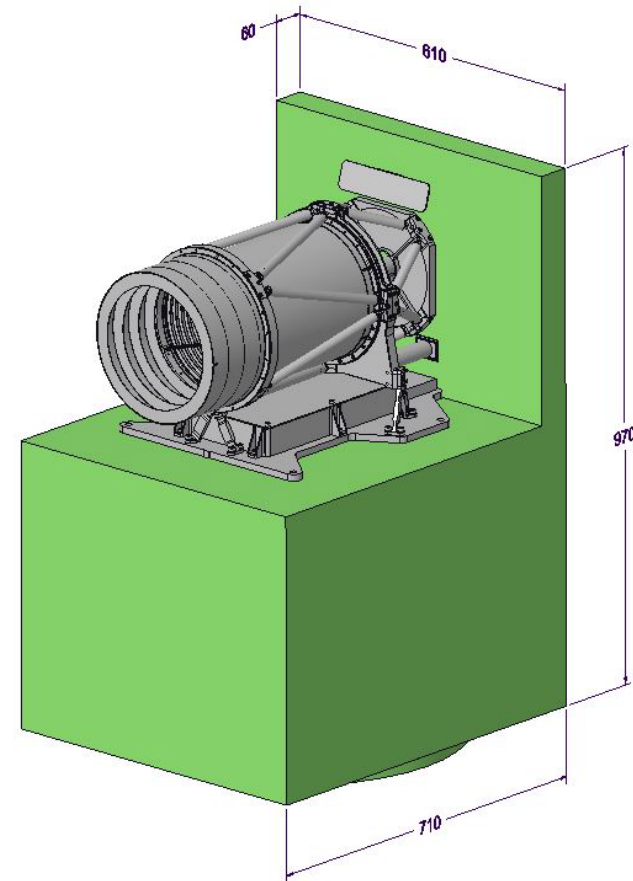
Wavelength coverage 250-1000 nm

Two channels:

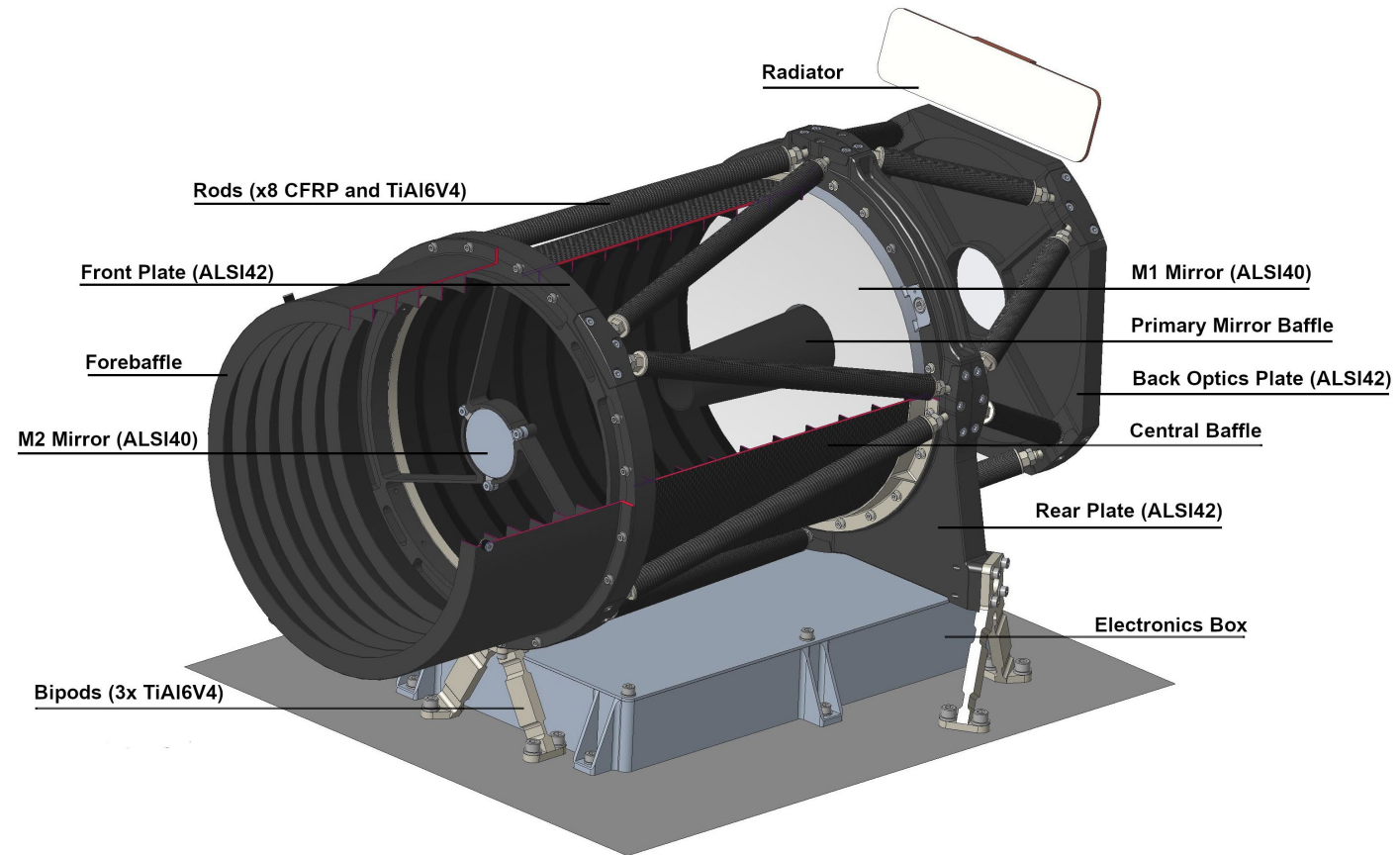
- VIS-NIR photometry
- NUV spectroscopy

Stable pointing

Expected mass 100-180 Kg

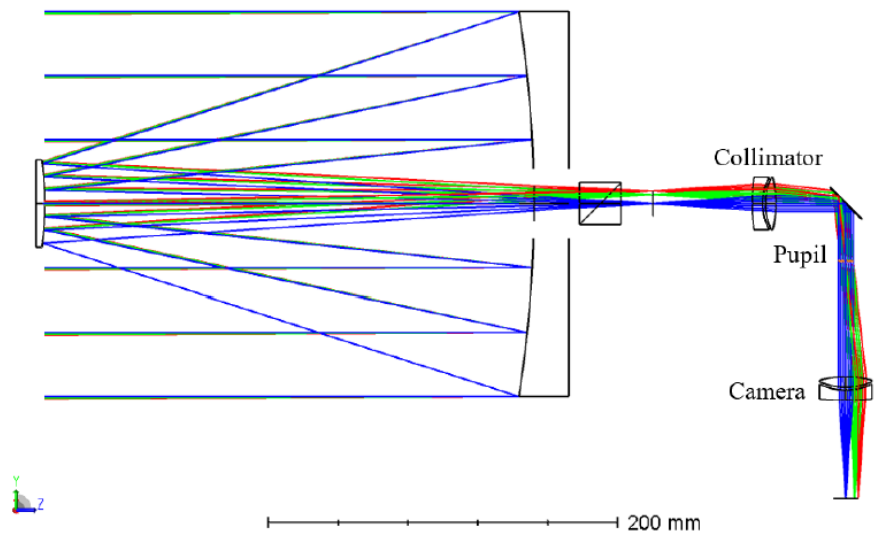


IACSAT-1: Optical bench

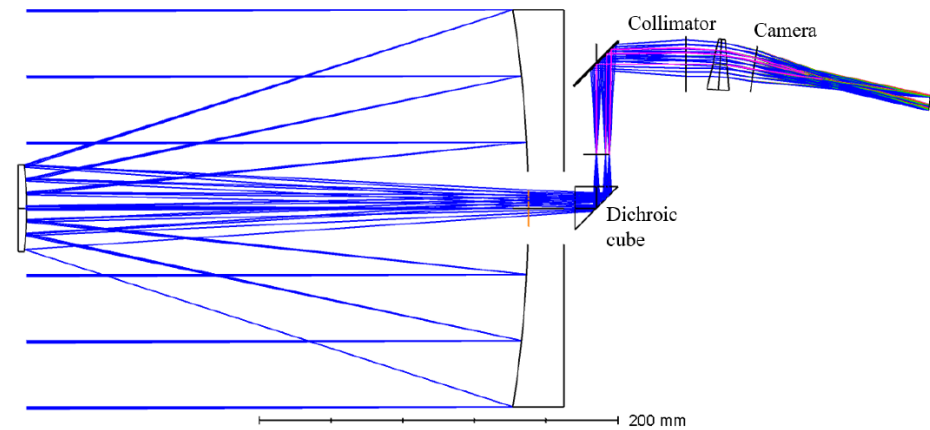


IACSAT-1: Two channels

VIS-NIR channel



NUV channel



IACSAT-1: Two channels

VIS-NIR channel

Broadband photometry

Wavelength range: 450-1000 nm

CCD 1024x1024

FOV 30'

- 1.6"/pixel resolution

Precision: 350 ppm

- 1 hour RMS, $V \sim 13.5$

NUV channel

Slitless spectroscopy

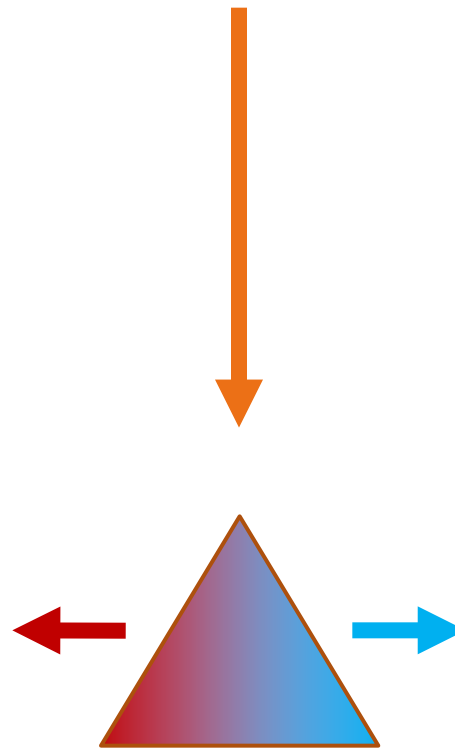
Wavelength range: 250-450 nm

CCD 1024x1024

Resolution element: 10 nm

SNR/pixel ~ 10

- 10 minutes exposure, $V_{\text{mag}} 14$



IACSAT-1: Orbit

600 km altitude

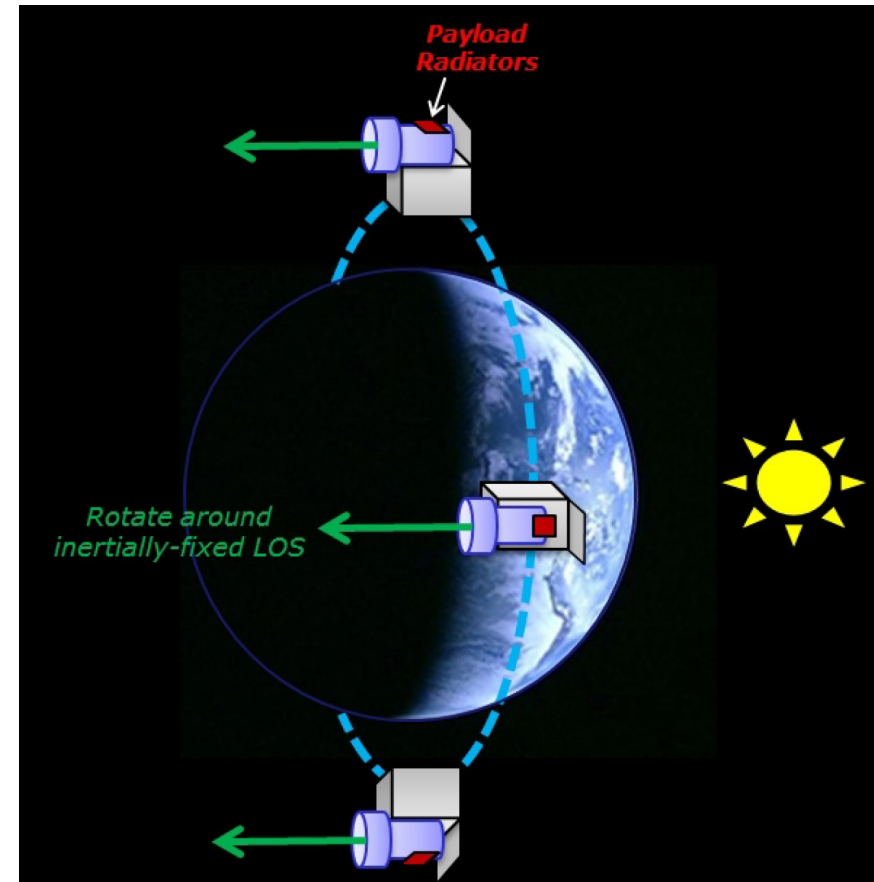
Sun-synchronous orbit

Riding the day-night terminator

- Limits impact of sunlight
- Limits impact of reflected light

Telescope rotation

- Optimal thermal stability
- Constraints to observing strategy



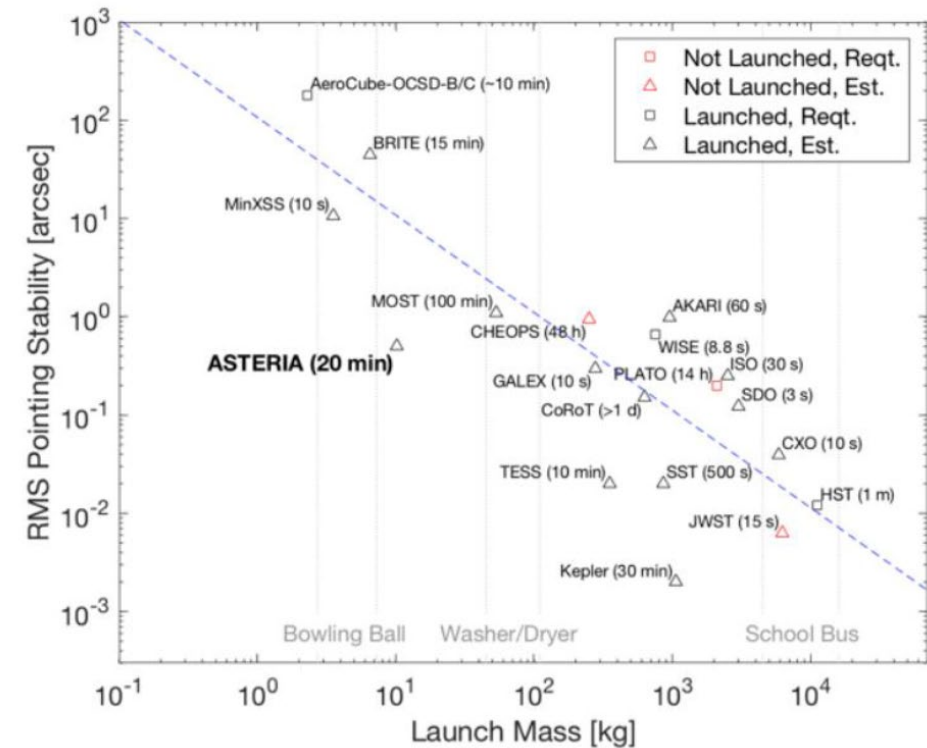
IACSAT-1: Pointing stability

2 arcsec RMS

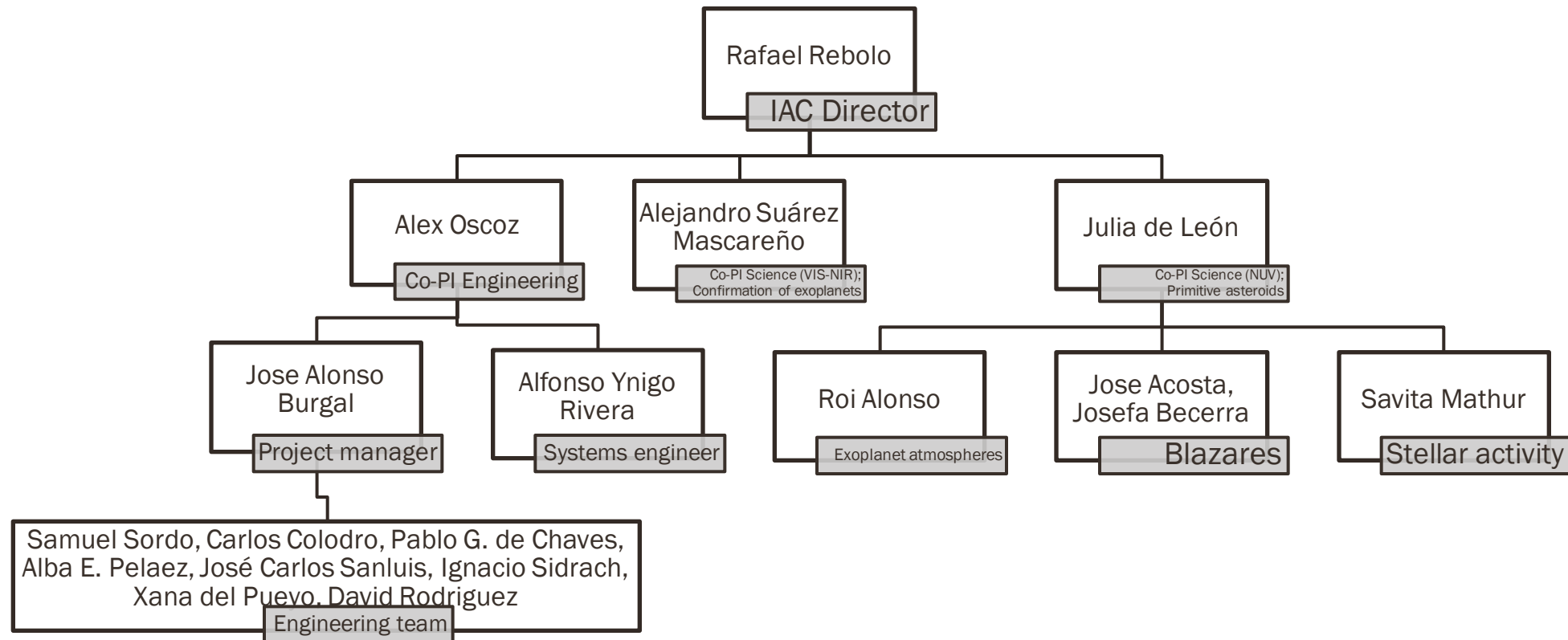
- Needed for high precision photometry
- Needed for spectroscopy
- Instrument on loop system

On par with:

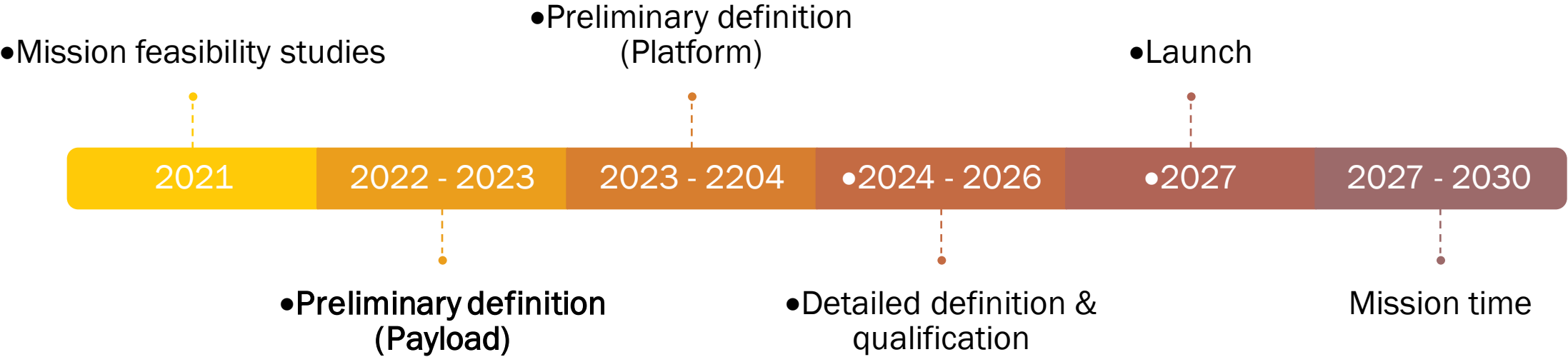
MOST
CHEOPS
ASTERIA
AKARI
WISE



IACSAT-1: Project organization



IACSAT-1: Project timescale



Summary

IACSAT-1: Multi-purpose space observatory

22 cm primary mirror

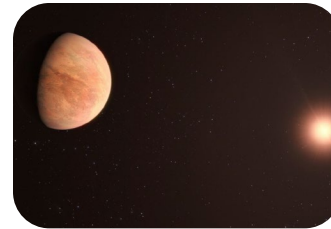
VIS-NIR photometry

NUV spectroscopy

Sun-synchronous orbit

Launch: 2027

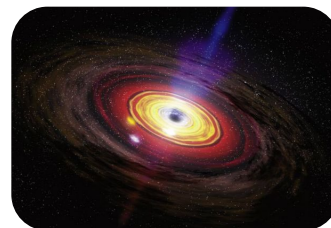
Currently in phase B



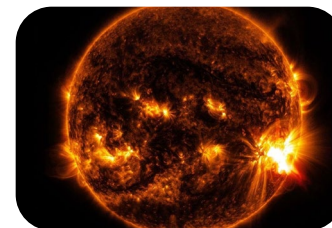
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